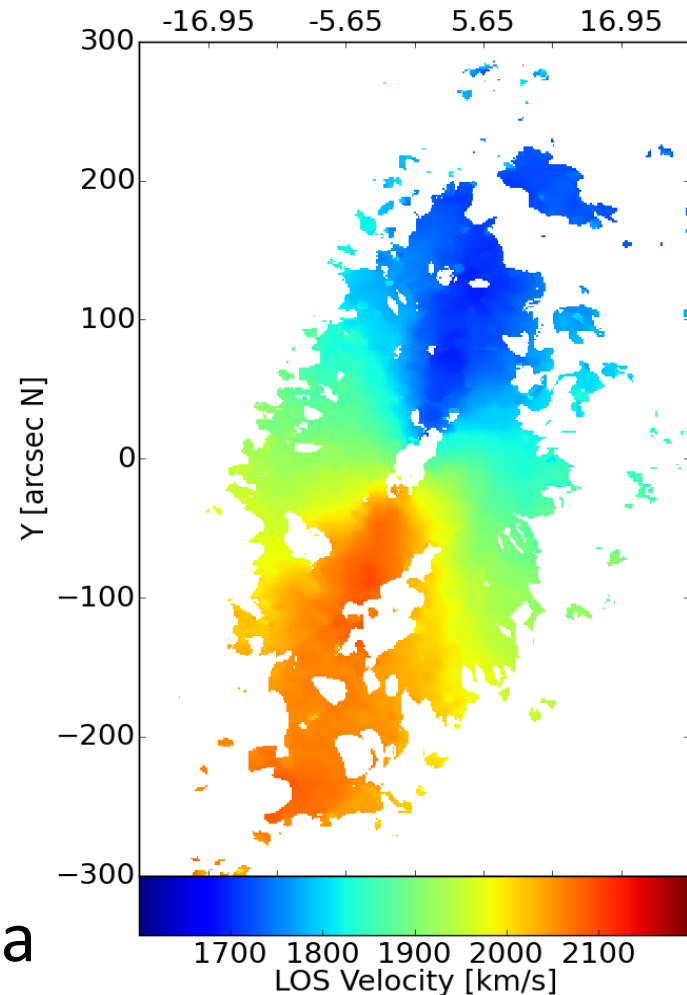


# Uncertainties in rotation curve estimation

J A Sellwood, Rutgers University

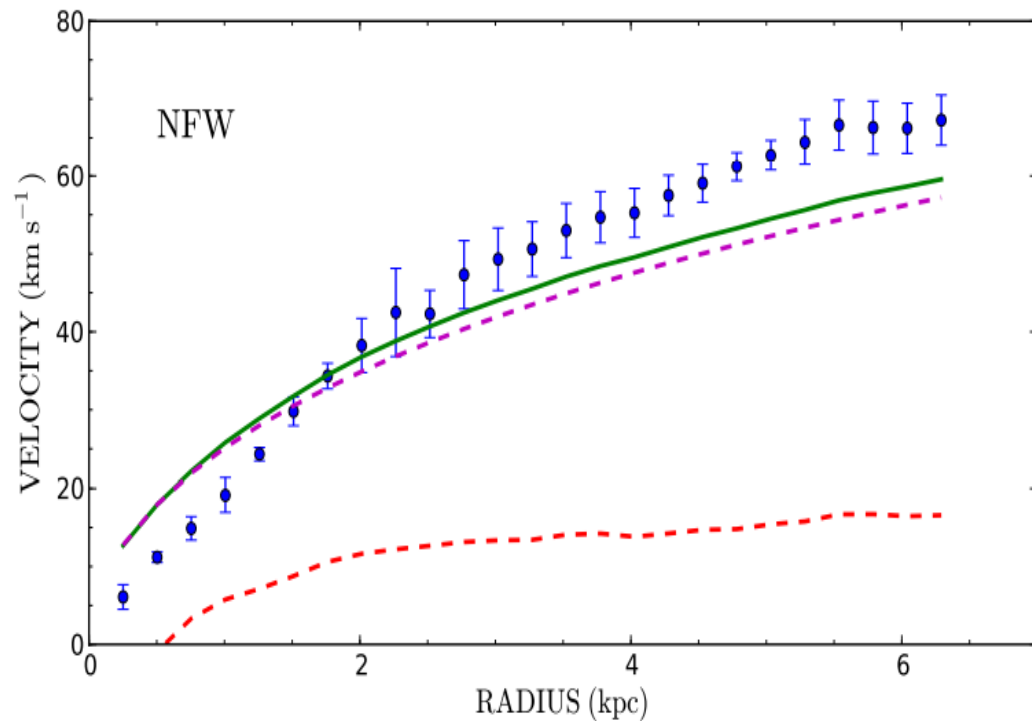
and Kristine Spekkens, RMC Canada

acknowledging significant advice from Tad Pryor  
and Carl Mitchell



# NGC 3109

Carignan et al. (2013)



- It is clear that NFW is a poor fit
- But we want to be able to **quantify the likelihood** that any given model is consistent with the data
  - can we make error estimates more **statistically robust?**

# Sources of systematic error

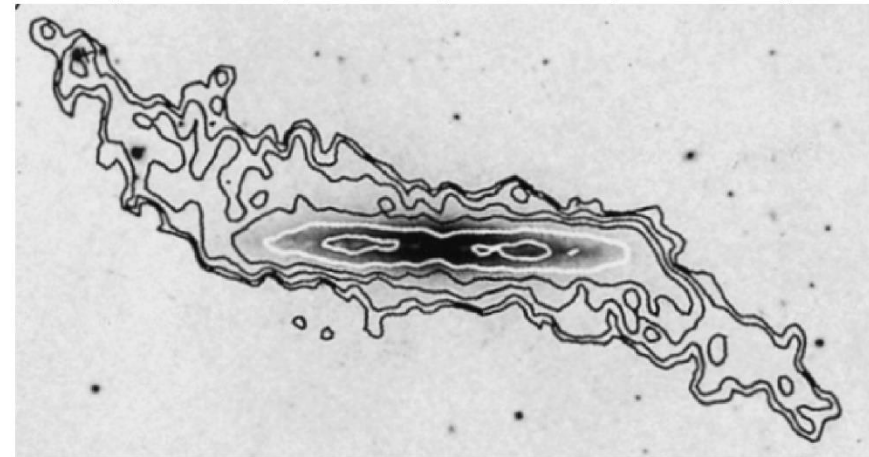
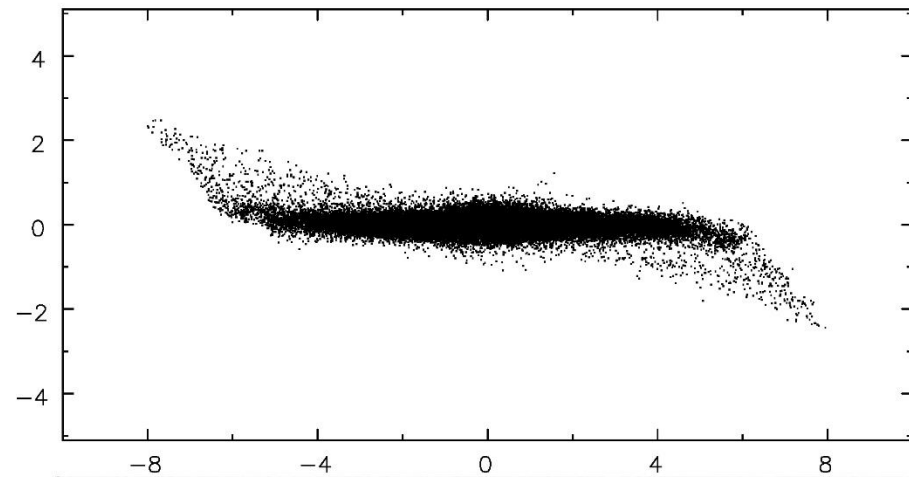
- A circular flow pattern in a flat disk may not fit a velocity map because of:
  1. bars and ovals (perhaps in the halo)
  2. warps
  3. spiral arm streaming
  4. asymmetries (lop-sidedness)
  5. turbulence
  6. out-of-plane motion – *e.g.* Sancisi's “beard”
  7. ...
- 1 & 2 can be modeled with *DiskFit* and tilted rings
- rest are best treated as sources of uncertainty
  - asymmetric drift corrections can be made once  $\langle V \rangle(R)$  is known

# Warps

(Model from Shen & JS 06 and NGC 4013 from Bottema 96)

- Theoretical prejudice:

- real warps are possible only in the low-density outer disk
- the massive part of a disk is quite stiff, and should be flat – unless the galaxy is disturbed



- Two conclusions

- Undesirable to use tilted ring fits that allow PA and  $i$  to vary in the inner disk
- also radial changes to the PA and  $i$  in a warp should perhaps be constrained to vary smoothly

# THINGS data for NGC 3621

(de Blok et al. 2008)

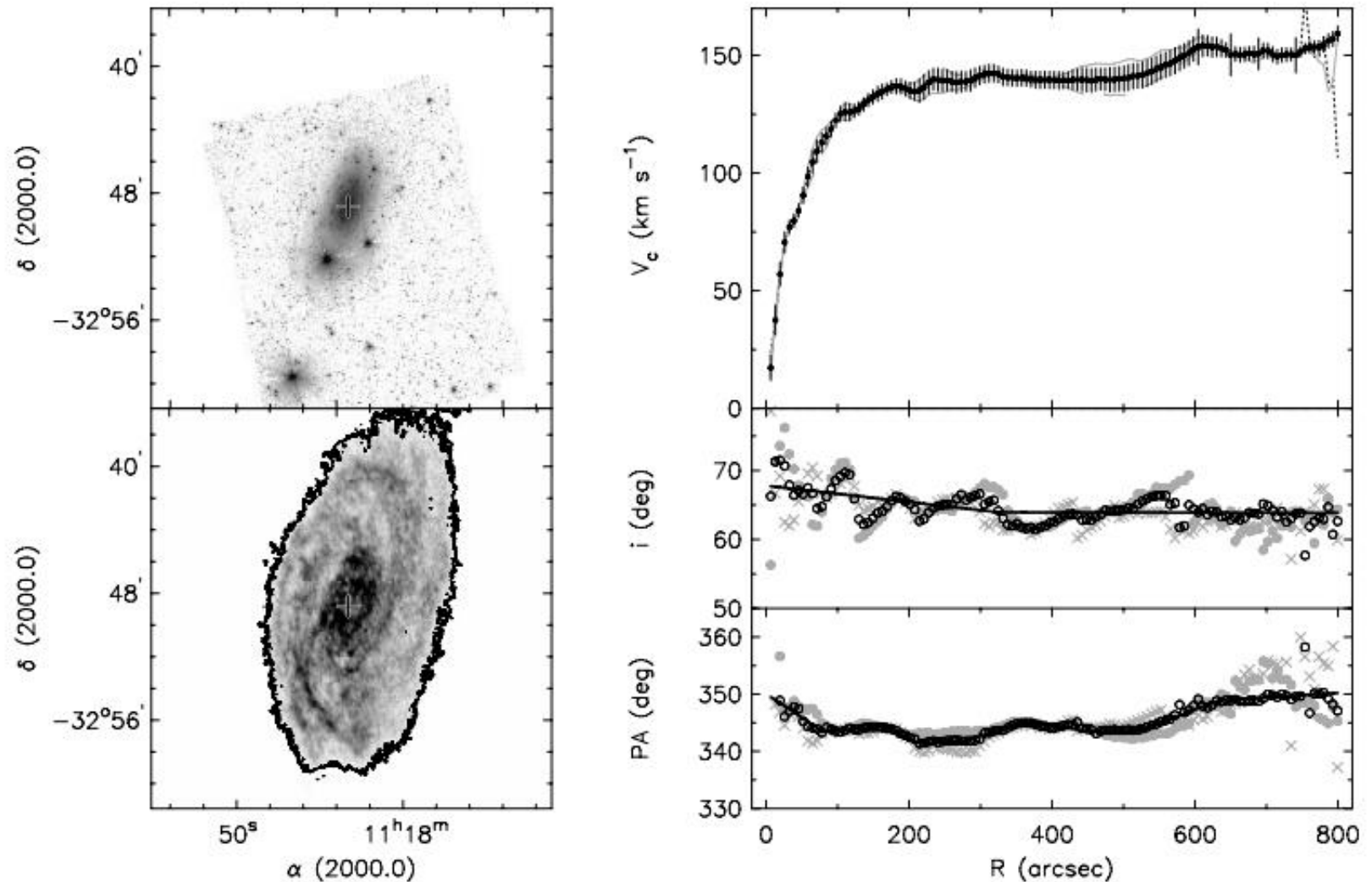
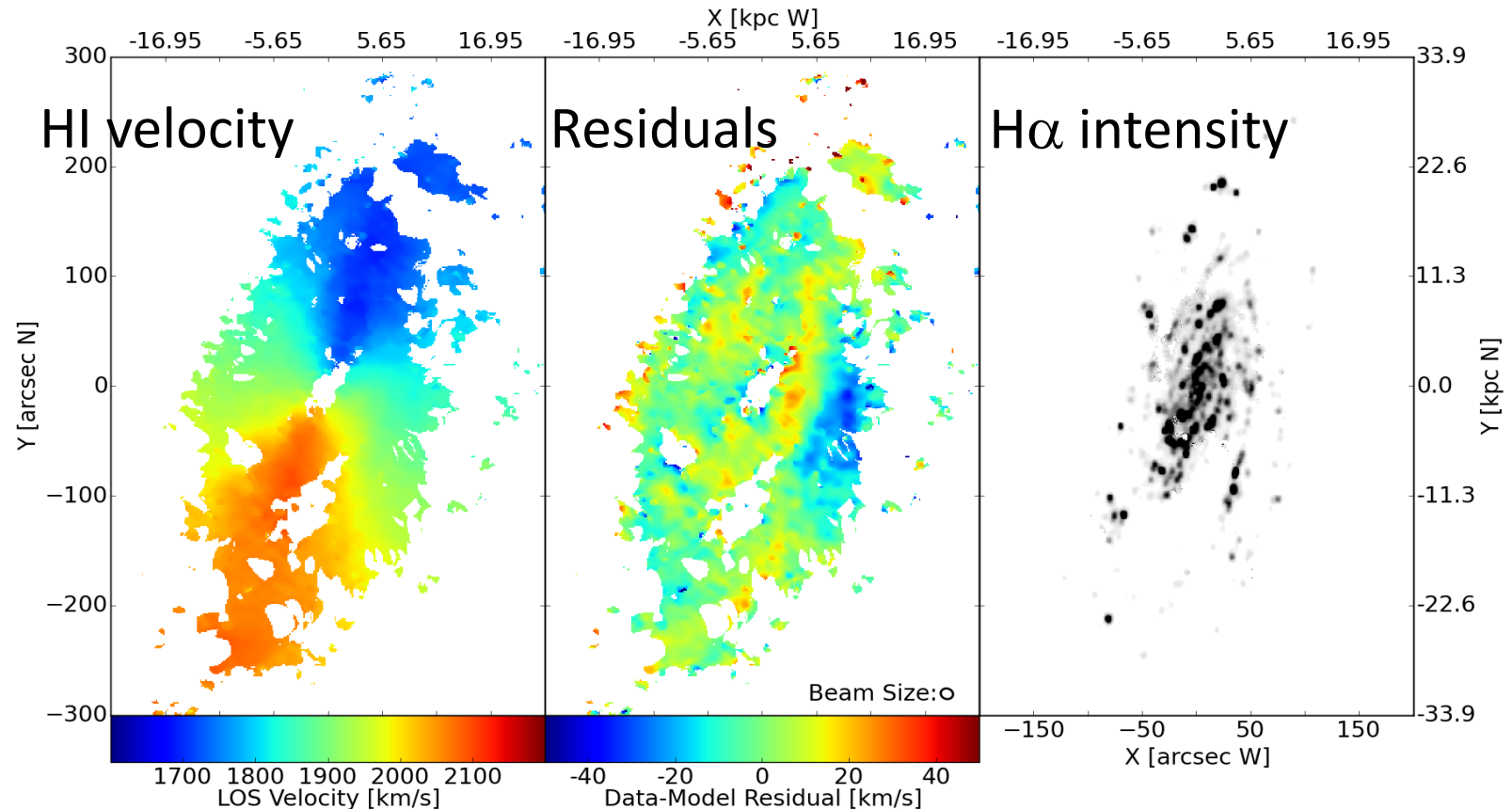


Figure 78. Summary panel for NGC 3621. See the [Appendix](#) for more information.

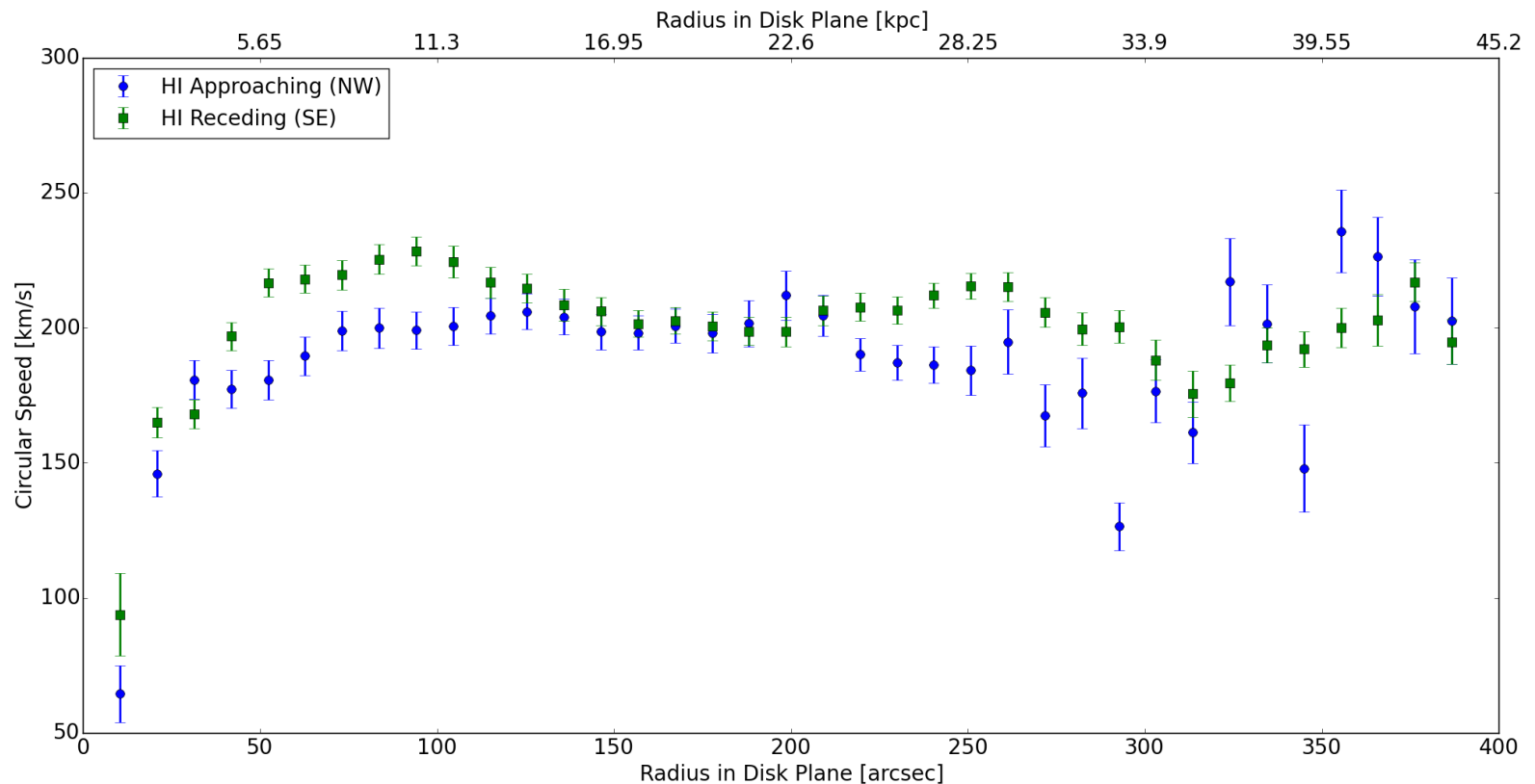
# Coherent residuals and asymmetries

- NGC 2280 data – Mitchell *et al.* (2015)
- Coherent residuals after fitting a flat disk with circular motion only – some correlation with spirals



# Fitted rotation curve

- Approaching and receding halves fitted separately
- Asymmetries not due to the clearest spirals, which are bi-symmetric



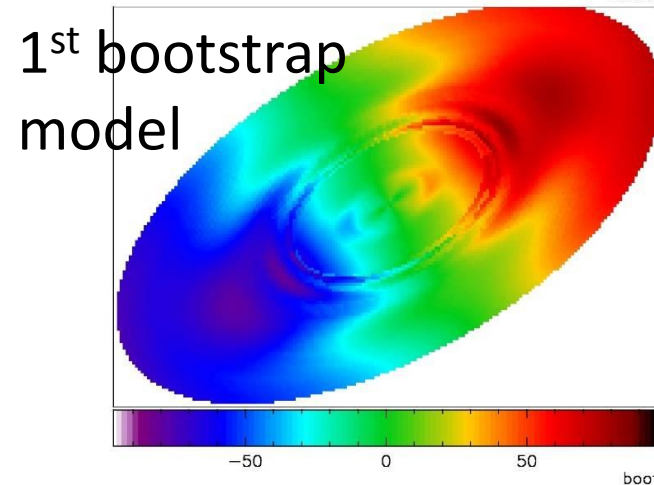
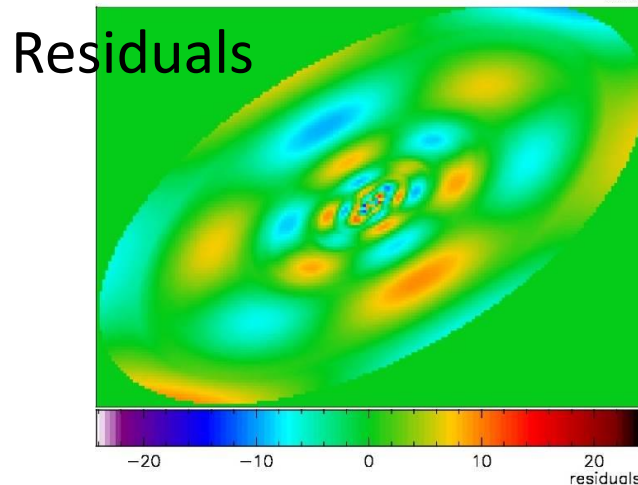
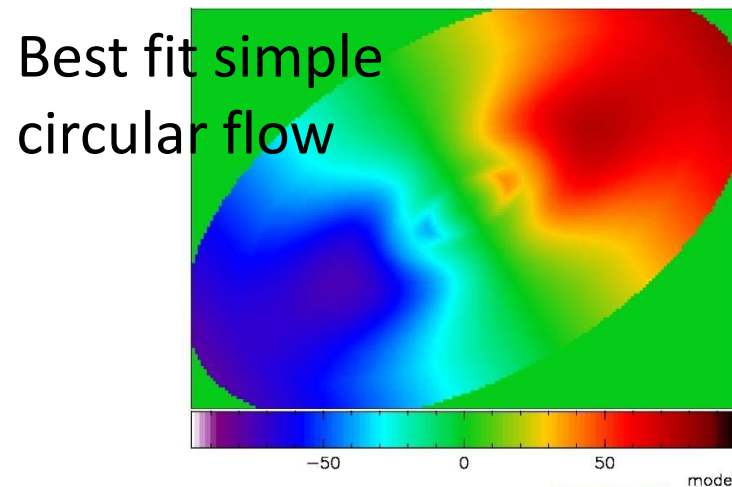
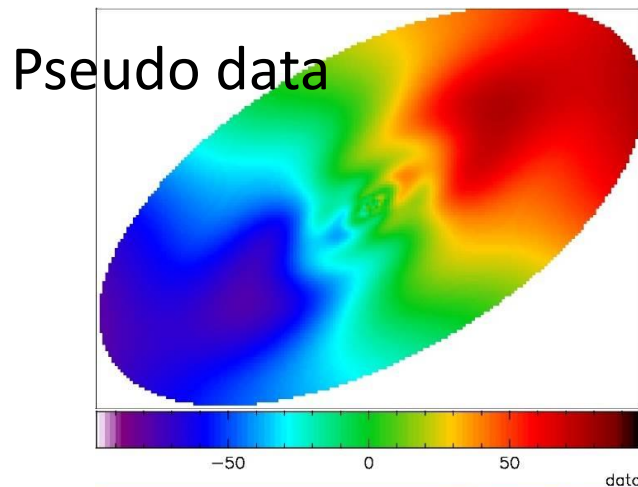
# Approaching/receding

- Traditional error in tilted ring studies is to add to the statistical error  $(V_{\text{app}} - V_{\text{rec}})/2$  in quadrature
  - no statistical basis for doing this, but
  - years of experience suggest it yields “realistic” errors
- *DiskFit*, on the other hand, uses bootstrap iterations
- Illustrate the idea with pseudo-data:
  - artificial map with no noise plus velocity distortions due to an  $m$ -armed spiral pattern
  - coherent residuals after fitting a flat, circular flow pattern



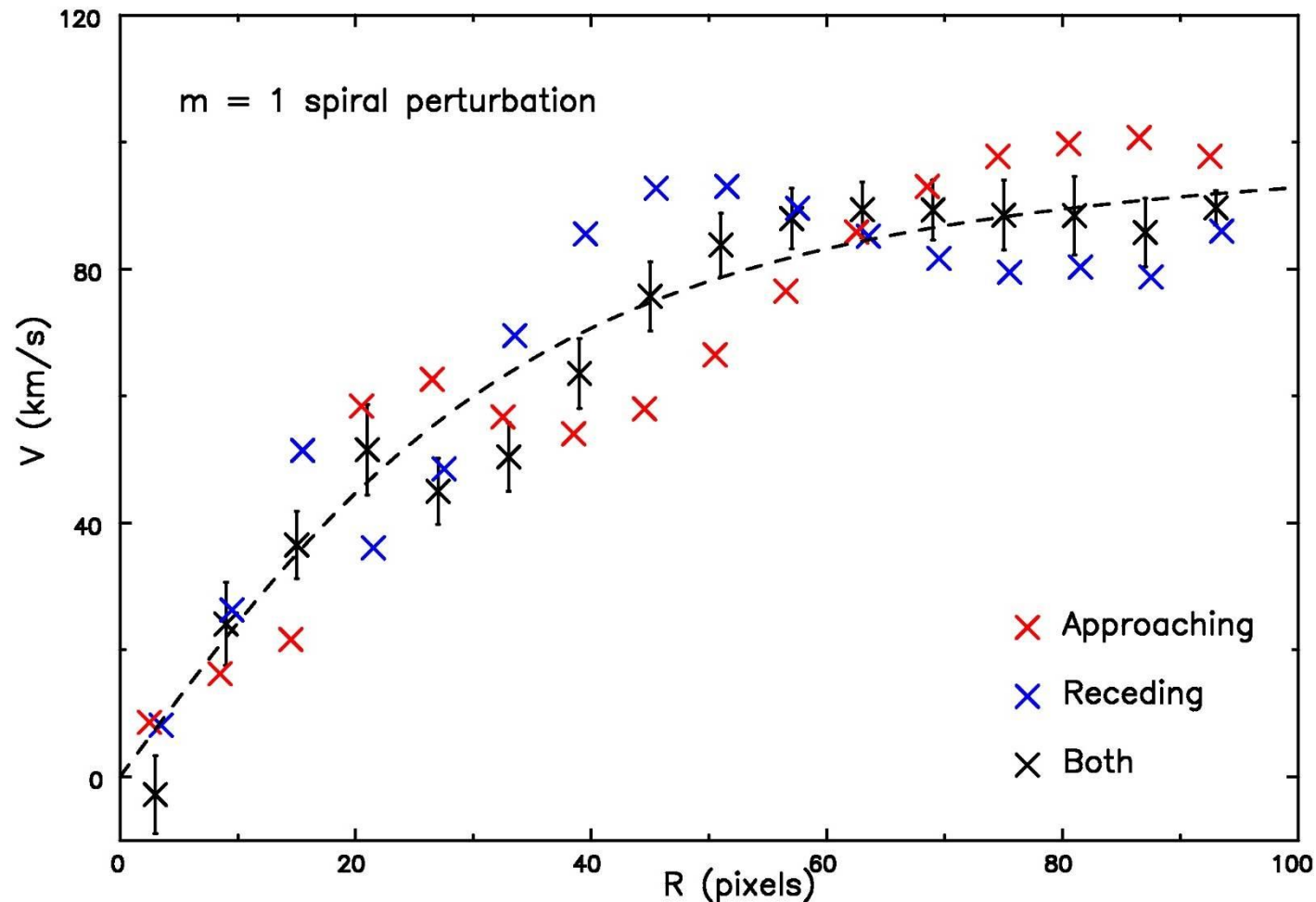
# Bootstrap estimation of uncertainties

- Modified bootstrap to take account of the coherent residuals
- Noise-free map with a spiral flow pattern added



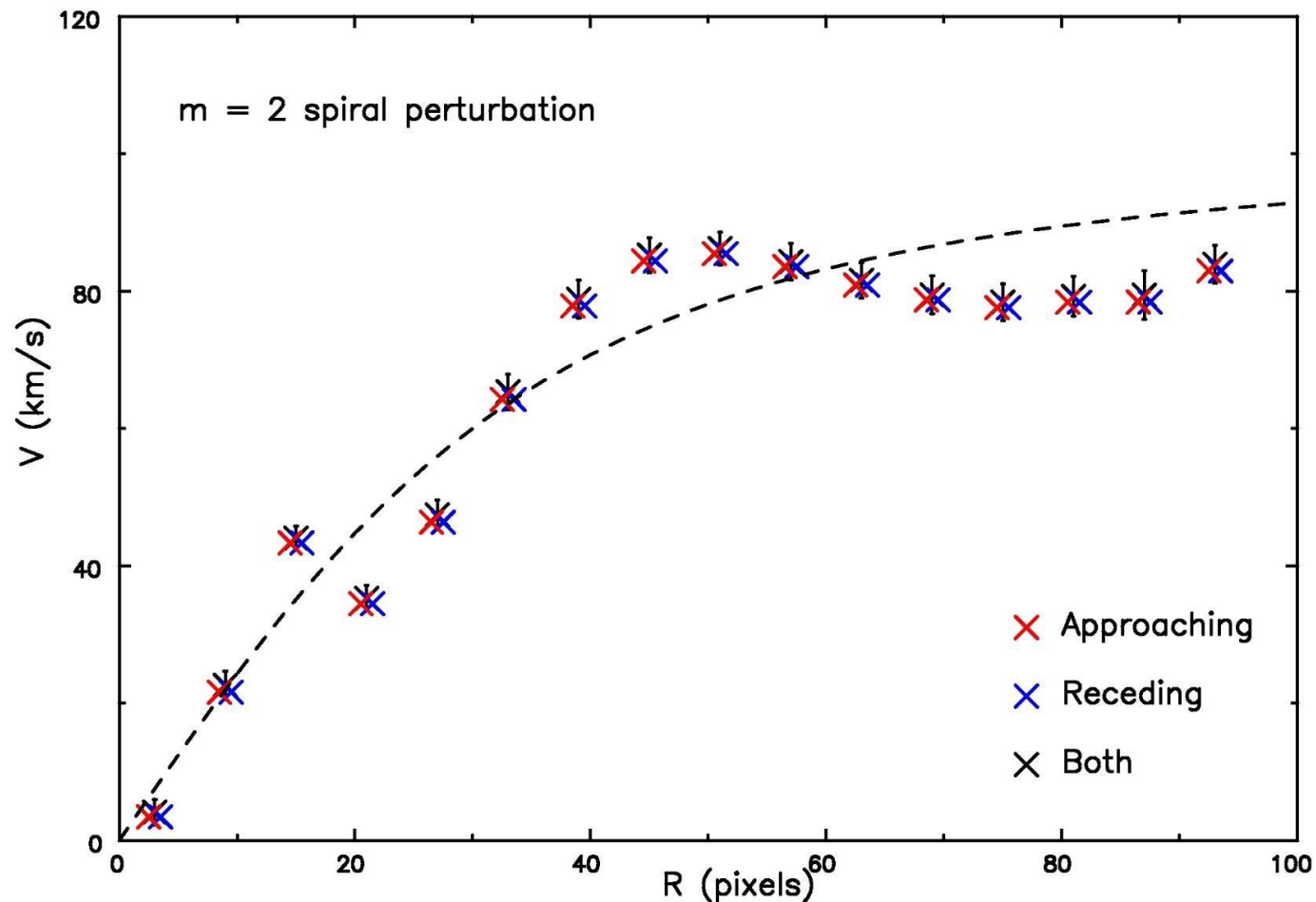
# Results

- An  $m=1$  spiral flow pattern added to the noise-free map
- Fitted with a flat, circular flow pattern with bootstrap errors



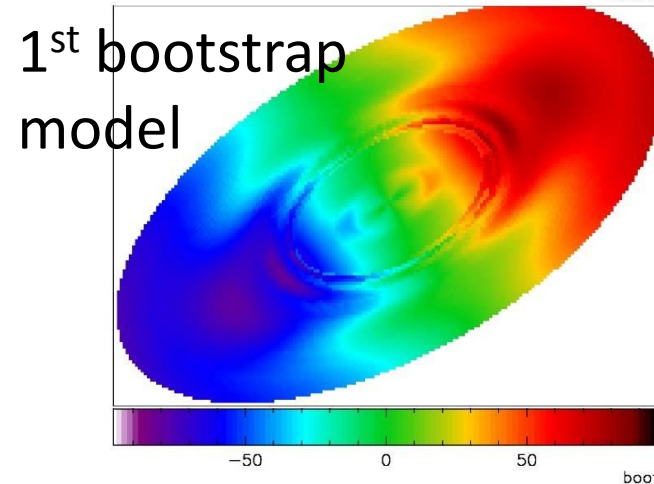
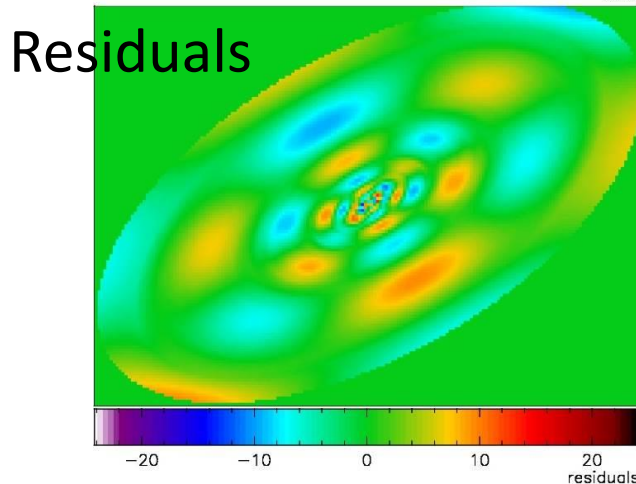
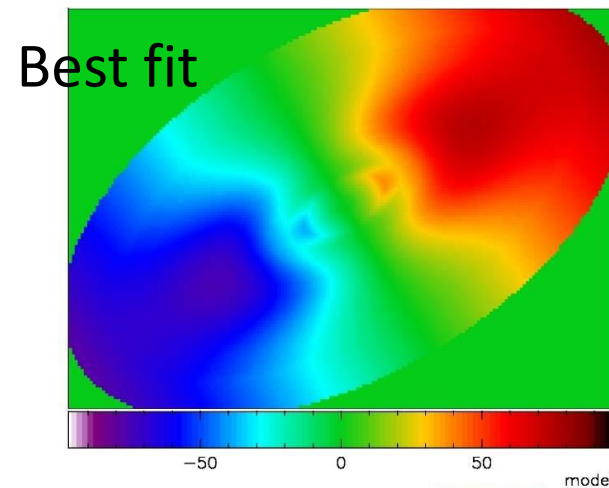
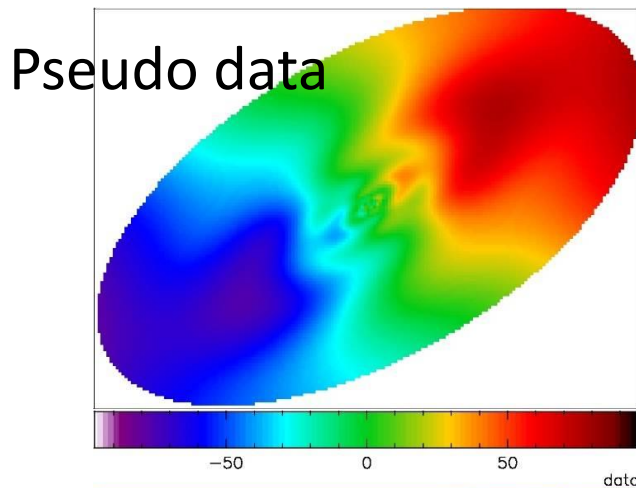
# Bi-symmetric spiral flow

- The systematic error in the velocity from an  $m=2$  spiral is the same on both sides
- Neither method for error estimation is adequate



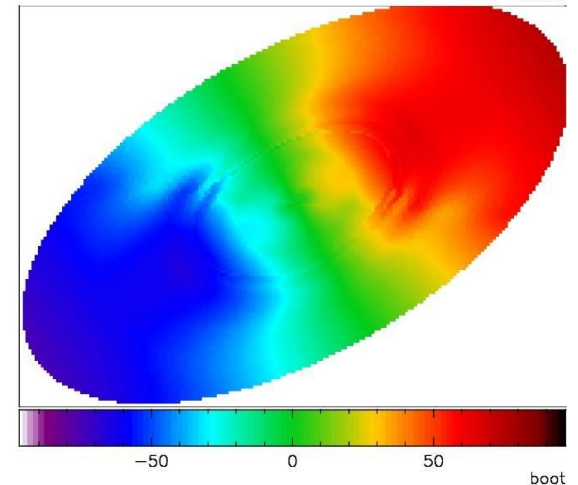
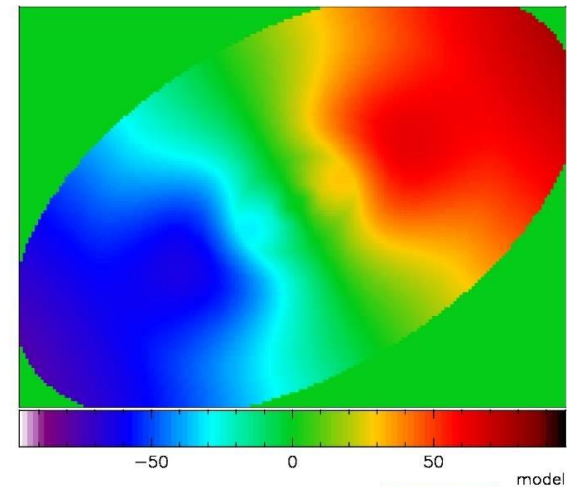
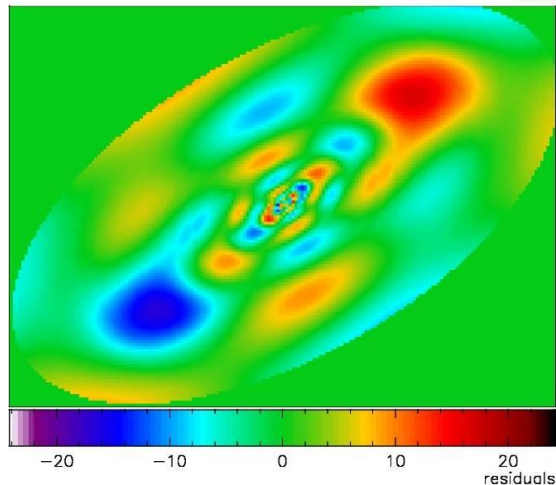
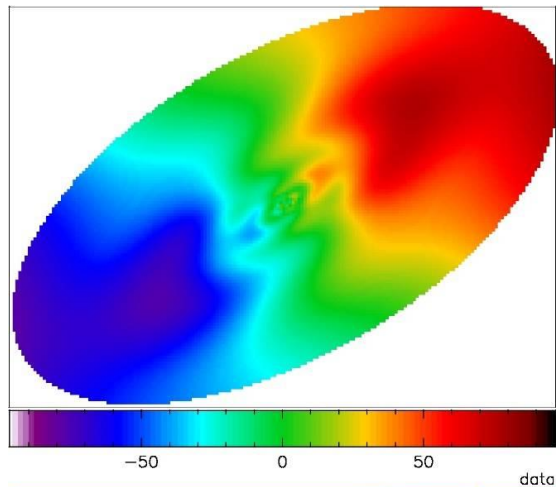
# Root of the problem

- Best fit model is too strongly affected by spiral arm crossings on the major axis
- Residuals not large enough to compensate



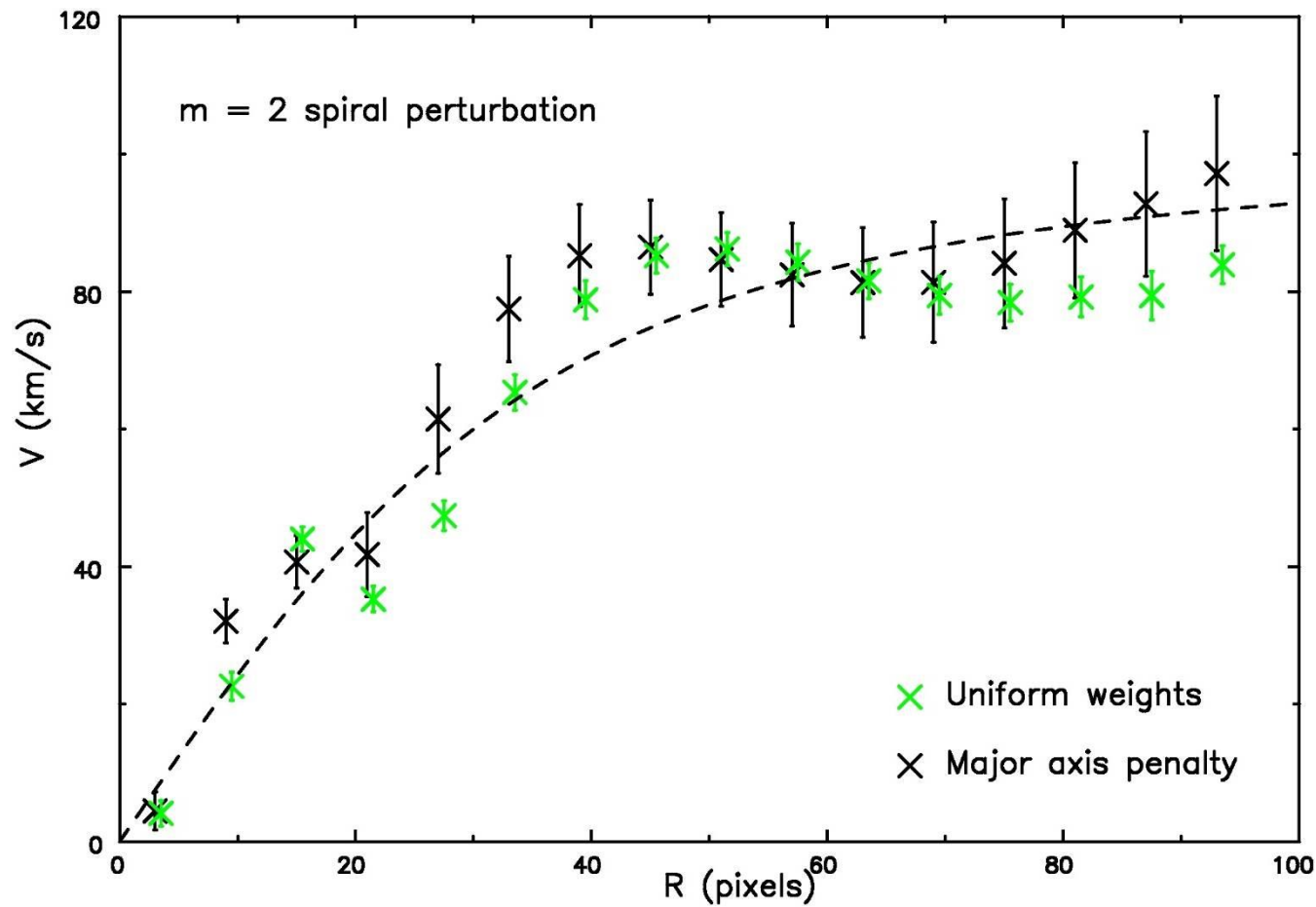
# We can do better

- Arm crossings on the major axis are given too much weight – so
- reduce the weight of points on the major axis



# Off-axis data influence fit more

- Best fit averages more over spiral phase
- Residuals, and therefore, bootstrap errors are increased



# Conclusions

- Massive inner disks should be flat
  - tilted ring fits there are undesirable
  - warps in the outer disk should be “smooth”
- Flat disk fits leave coherent patterns of residuals due to spirals, asymmetries, turbulence, *etc.*
- Need a statistically valid means to estimate these systematic uncertainties in  $V_c(R)$ 
  - $(V_{\text{app}} - V_{\text{rec}})/2$  not statistically robust and **does not take account of errors due to 2-arm spiral flows**
  - modified bootstrap is one possible way to do it
  - ideas still “in progress”, and need extensive testing
- Errors must be estimated properly in order to mount a quantitative challenge to theoretical models